Interstellar Travel

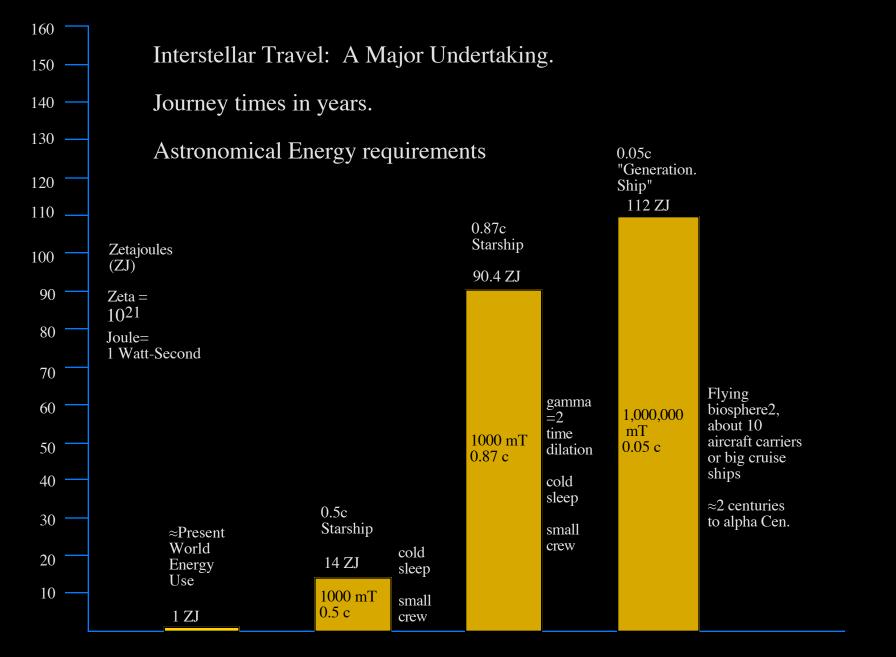
Why Interstellar Travel

Where Can We Go?

How Can We Get There

What Would an Interstellar Civilization Might Be Like

G. David Nordley OryCon 2010



Why Interstellar Travel?

Provide realtime human guidance to exploration robots.

Entertainment value for actual and vicarious participants

Commerce? Unique artifacts?

- Extrasolar artwork by aliens or human colonists?
- Archeological SETI exhibits for Earth museums

Not for...

Resources--much cheaper to mine or make in the solar system



Extrasolar Colonies for More Living Room?

Short term:

--Isolation.

Some folks just can't get along

... or are tired of putting up with idiots.

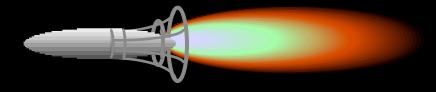
--Empire building

No room on Earth for this antisocial occupation, but go grow your own?

Long term:

--External disasters we can't cope with otherwise? (When Worlds Collide) we know Jupiter-sized rogue planets exist...

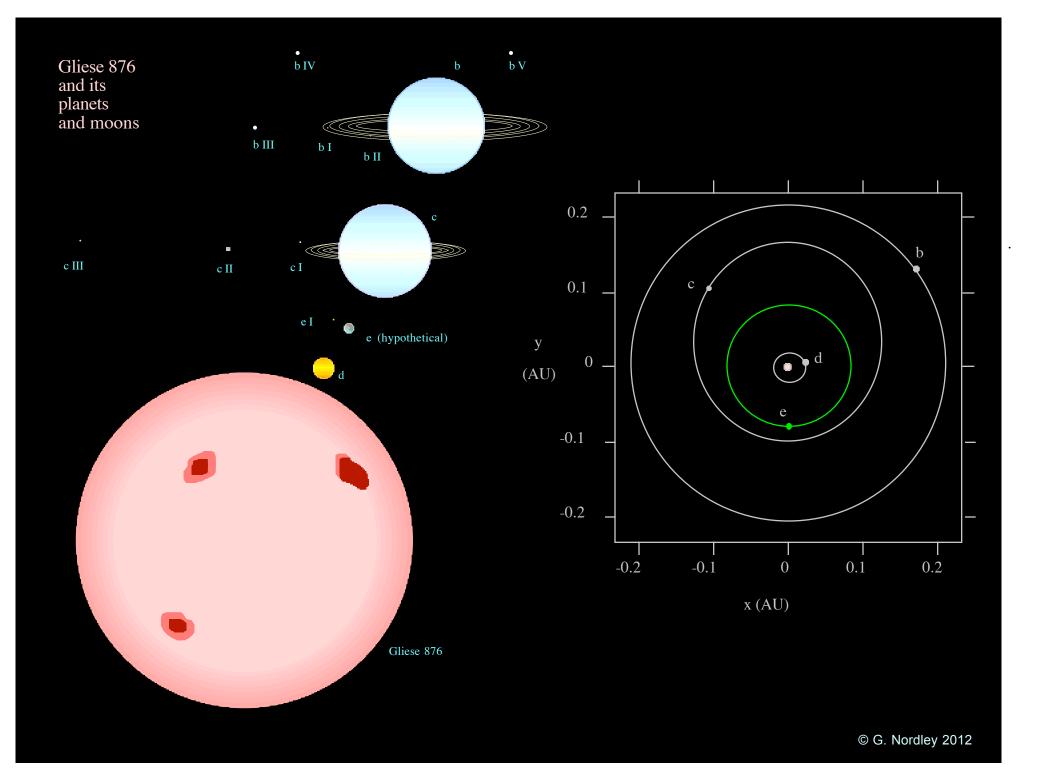
--Our habitable zone is moving out as the Sun ages and expands;
we might buy a few billion years with sunshields or moving Earth...
BUT red dwarfs last for around a trillion years.

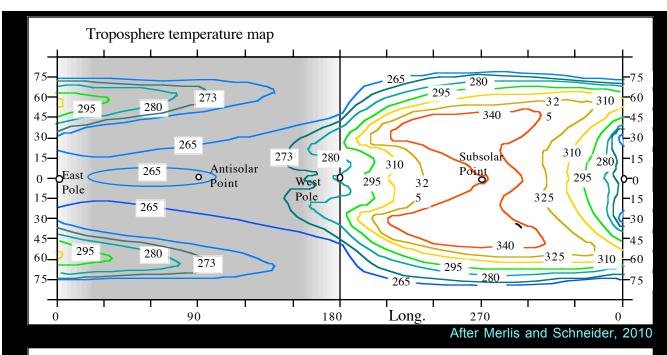


Nearby Star Destinations

Name Sol	Spectrum G2V	Luminosity 1	Distance 0	"AU" 1	"year" 1	Comments Ultraviolet an issue <u> </u>
Proxima Centauri	M5.5Ve	0.0007	4.223	0.026	0.014	Flares.
alpha Centuari A	G2V	1.55	4.394	1.25	1.32	Ultraviolet an issue.
alpha Centuari B	K1V	0.528	4.394	0.727	0.672	olliaviolet all issue.
Barnards Star	M4Ve	0.003	5.938	0.057	0.034	Flares.
Wolf 359	M5Ve	0.0002	7.674	0.014	0.007	Flares.
Lalande 21185	M2V	0.026	8.309	0.162	0.114	Flares.
UV Ceti A	M5V	0.001	8.561	0.031	0.017	Flares.
UV Ceti B &C?	M5Ve	0.0006	8.561	0.024	0.013	Flares.
Sirius A	A1V	23.5	8.601	4.84	7.2	Stable orbits difficult, deadly UV
Sirius B	B9VII	0.003	8.601	0.057	0.035	White Dwarf Deadly UV.
Ross 154	M3.5Ve	0.002	9.693	0.046	0.027	Flares.
Ross 248	M5Ve	0.001	10.35	0.033	0.018	Flares.
epsilon Eridani	K2V	0.346	10.5	0.588	0.516	Young: flares, radioactivity, eccent. Jovian 🦳
Luyten 789-6	M5.5Ve	0.002	10.87	0.04	0.023	Flares.
Ross 128	M4Vfl	0.003	10.89	0.051	0.030	Flares.
61Cygni A	K5V	0.137	11.36	0.37	0.289	
61Cygni B	K7V	0.085	11.43	0.291	0.214	
Procyon A	F5V-IV	7.17	11.41	2.68	3.43	Stable orbits difficult, deadly UV
Procyon B	A5VII	0.0007	11.41	0.026	0.014	White Dwarf. Deadly UV.
Gliese 725B	M3.5V	0.009	11.47	0.097	0.063	
Gliese 725A	M4V	0.022	11.64	0.148	0.102	
Groombridge 34A	M1V	0.024	11.64	0.154	0.107	
Groombridge 34B	M3.5Ve	0.003	11.64	0.053	0.032	Flares
Lacaille 9352	M0.5V	0.042	11.69	0.204	0.148	
epsilon Indi	K4.5V	0.221	11.83	0.47	0.390	
tau Ceti	G8V	0.466	11.9	0.682	0.621	
Luyten 725-32	M4.5Ve	0.002	12.12	0.048	0.029	Flares.
Luytens Star	M3.5V	0.009	12.39	0.095	0.062	
Kapteyns Star	M1VI	0.013	12.78	0.116	0.077	
Lacaille 8760	K7Vfl	0.059	12.87	0.243	0.175	
Kruger60A	M3V	0.012	13.07	0.109	0.073	Flares.
Kruger60B	M3V	0.006	13.07	0.079	0.050	Flares.
Gliese 876	M2V	0.012	15.33	0.109	0.064	Hot Neptune, 2 "warm" Jupiters 🛛 👝 🥏
Groombridge 1618	K5V	0.074	15.89	0.271	0.196	
Luyten 354-89	M1.5V	0.029	16.11	0.172	0.121	
70 Ophiuchi A	KOV	0.495	16.39	0.703	0.645	
70 Ophiuchi B	K4V	0.118	16.39	0.343	0.263	
40 Eridani A	K1V	0.434	16.45	0.659	0.595	
40 Eridani B	A7VII	0.003	16.45	0.056	0.034	White Dwarf Deadly UV
40 Eridani C	M5.5V	0.006	16.45	0.08	0.051	
						© G. Nordley 2012

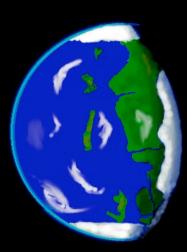
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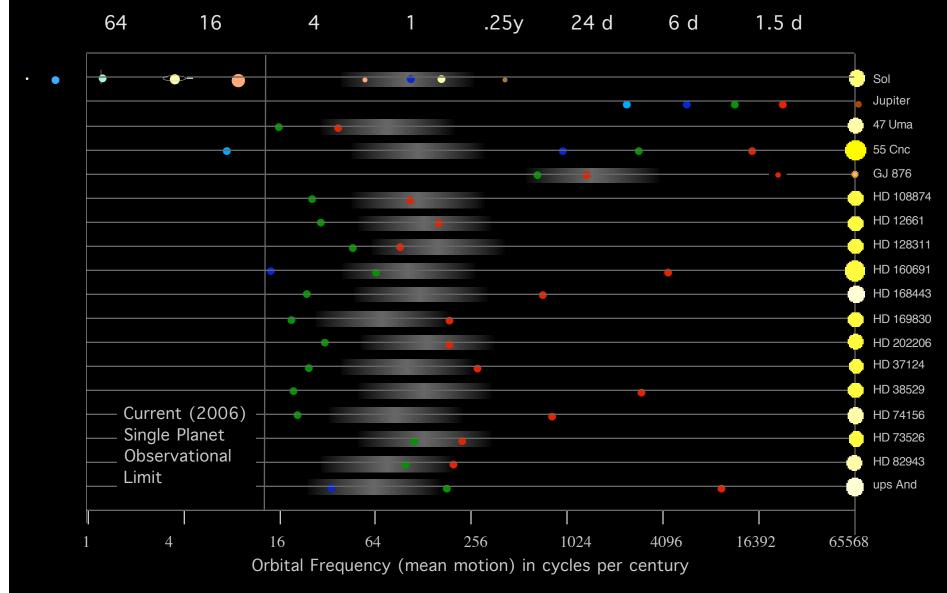


Surface concept subduction zones Mtns 75--75 Plateau, e, 60--60 dark sea • Mtns 45olcan -45 0 os 30--30 15--15 East Subsolar Ice Point 0 Pole ō **Ò-**0 0-¢ West Cap Pole 15--15 30--30 Volcanos 45--45 1 dark sea 60--60 75--75 90 270 180 0 0 Long.

TIDE LOCKED HABITABLE PLANETS

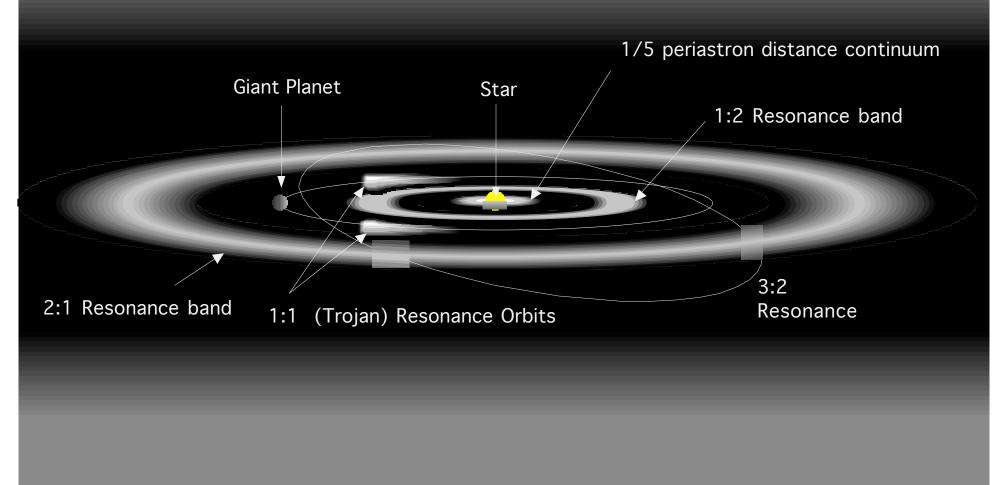


Orbital Period Chart: Sol, Jupiter, and Some Extrasolar Systems

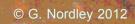


Resonances can allow stable orbits near giant planets

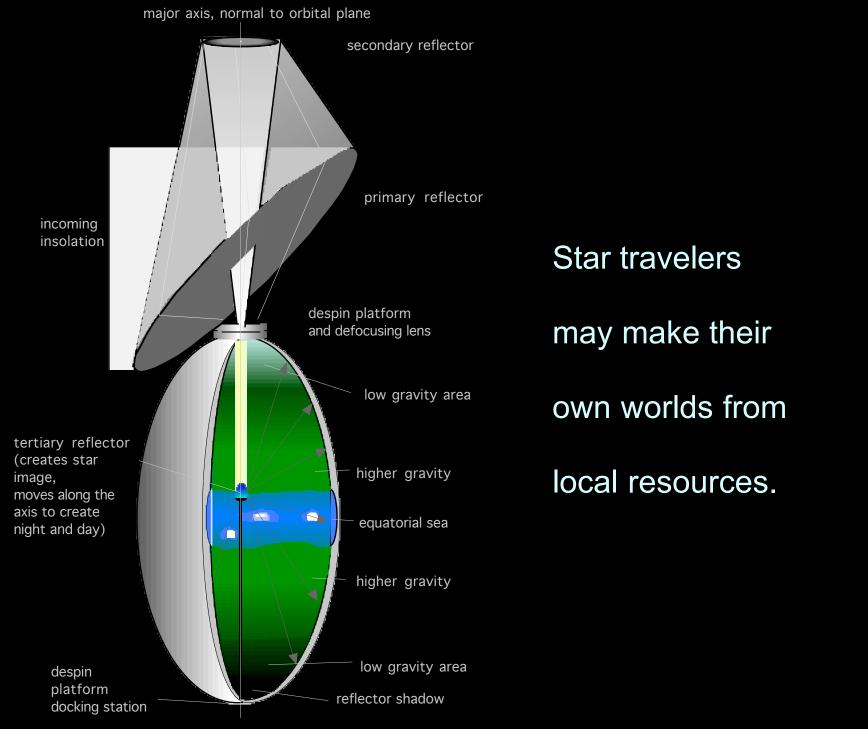




Habitable or Terraformable Moons?



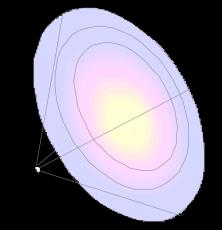




HOW DO WE GET THERE?



--SAILS



Interstellar Flight: Energy, Velocity, Time

$$KE = m c^2 (\gamma - 1) \qquad \gamma = \sqrt{(1/(1-v^2/c^2))}$$

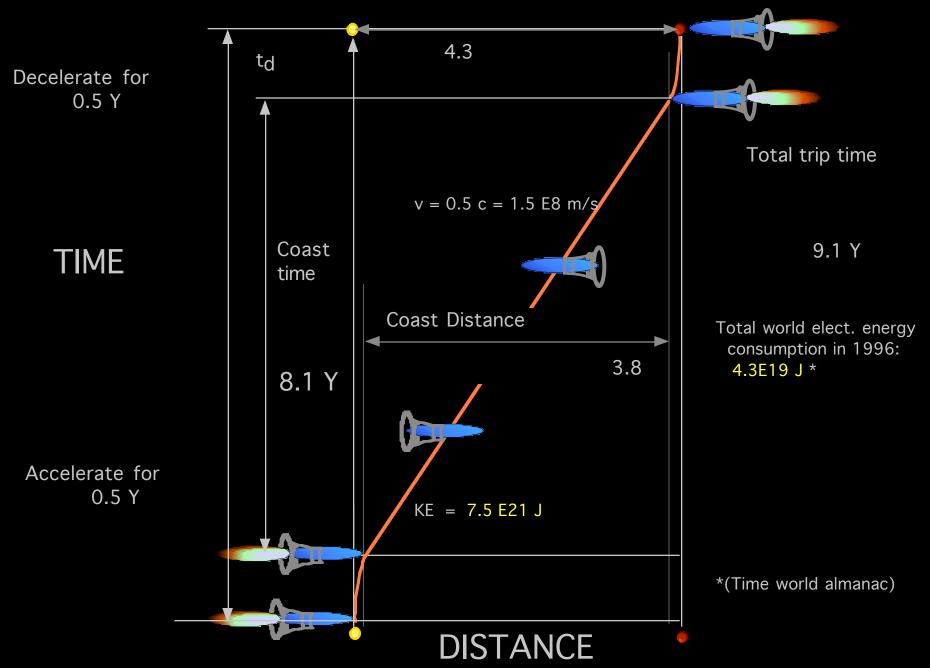
KE = Kinetic Energy γ = Lorentz factor m = mass v= velocity

For constant acceleration a in the spaceship frame

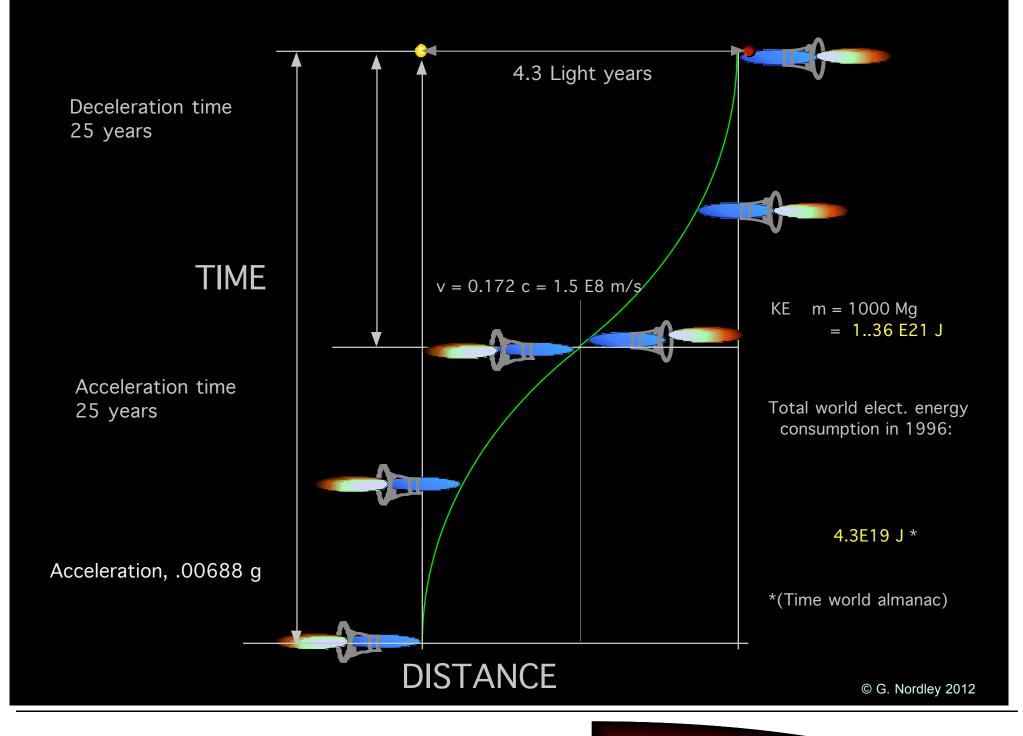
distance $s = c^2(\gamma - 1)/a = KE/(m^*a)$

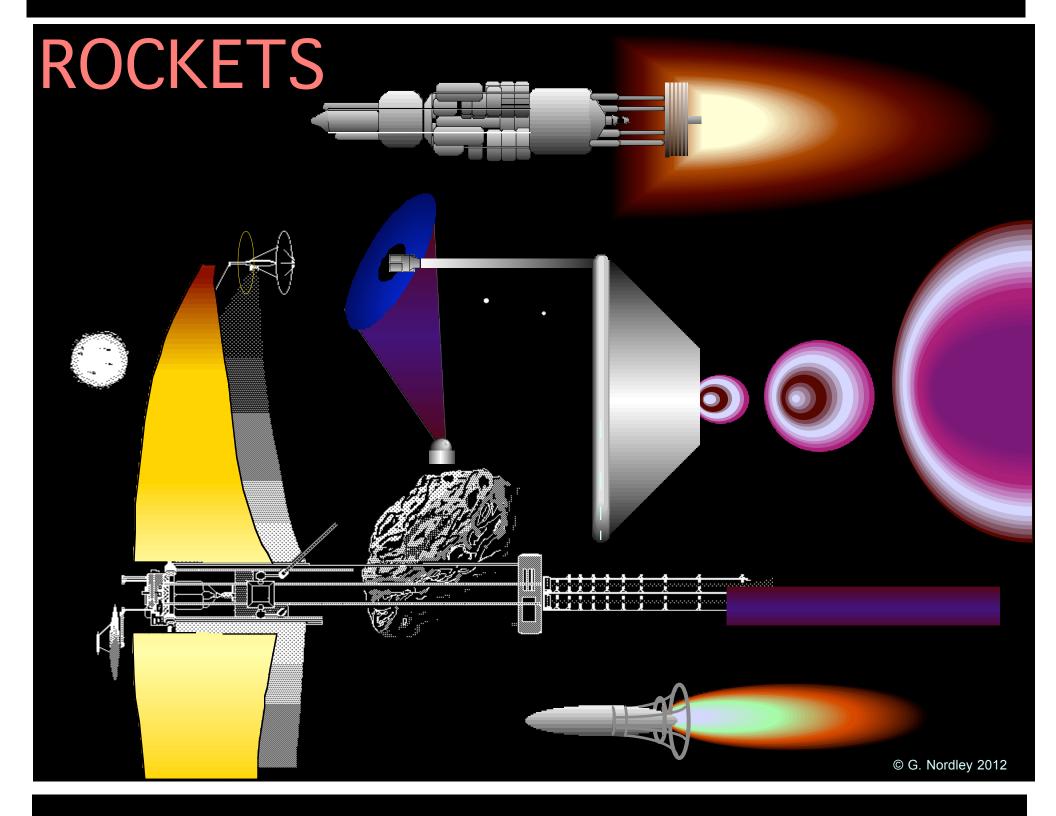
acceleration time $\Delta t = v/a + v^*s/c^2$

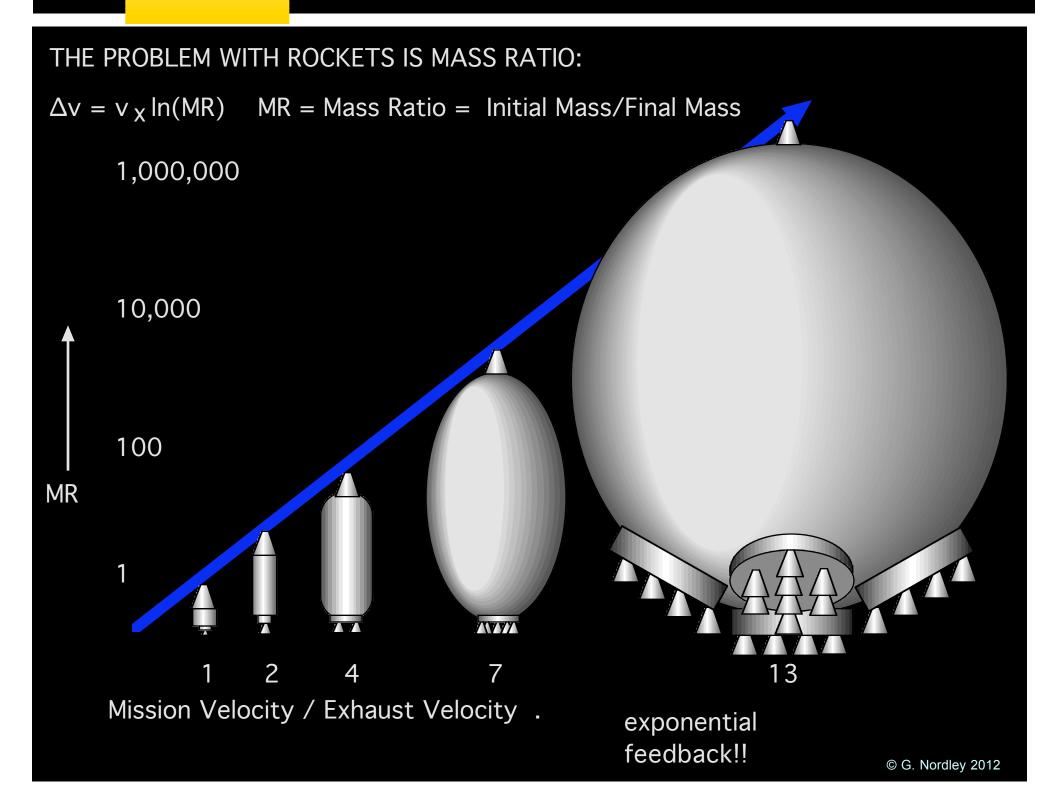
Interstellar Flight: Accelerate, Coast, Decelerate



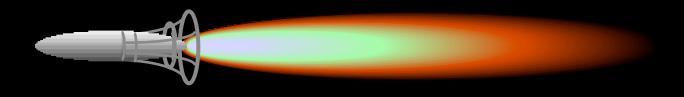
Interstellar Flight: Minimum Acceleration







Magnetically Confined Continuous Fusion



Well Studied for power production

Potential Exhaust Velocities up to \approx 0.1 c,

High mass ratio means ≈ 0.15 c cruise velocity

"Magnetically confined fusion is about twenty years in the future..." (...and always will be?)



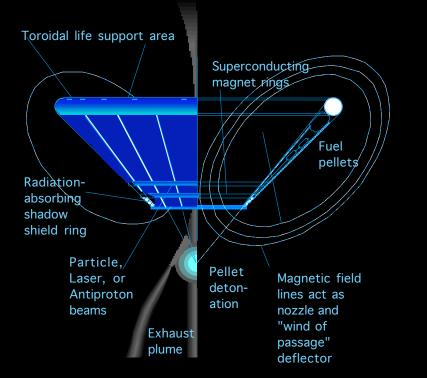
Exhaust velocities \approx 0.01-0.1c

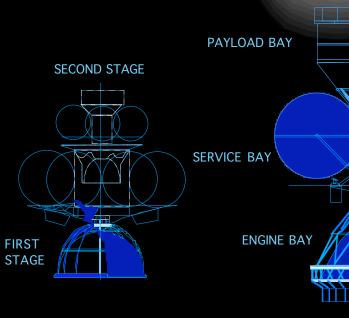
Alpha Centauri trip in decades

May use to decelerate

Still working on pellet detonation....

Micropulse fusion vehicle (after Hyde)



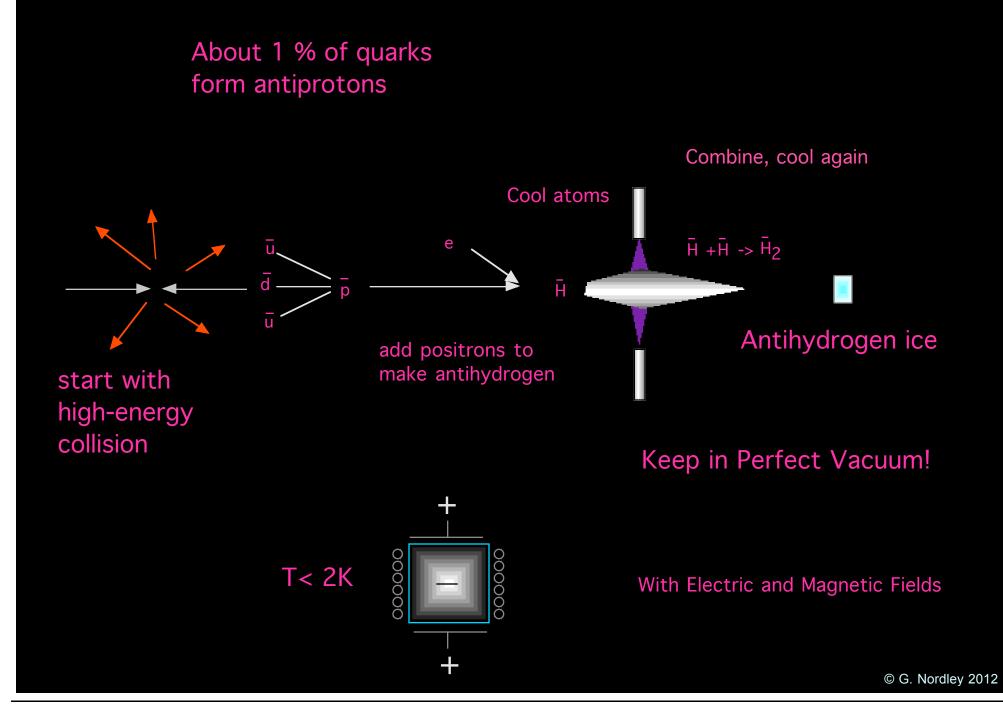


THE BRITISH INTERPLANETARY SOCIETY PROBE

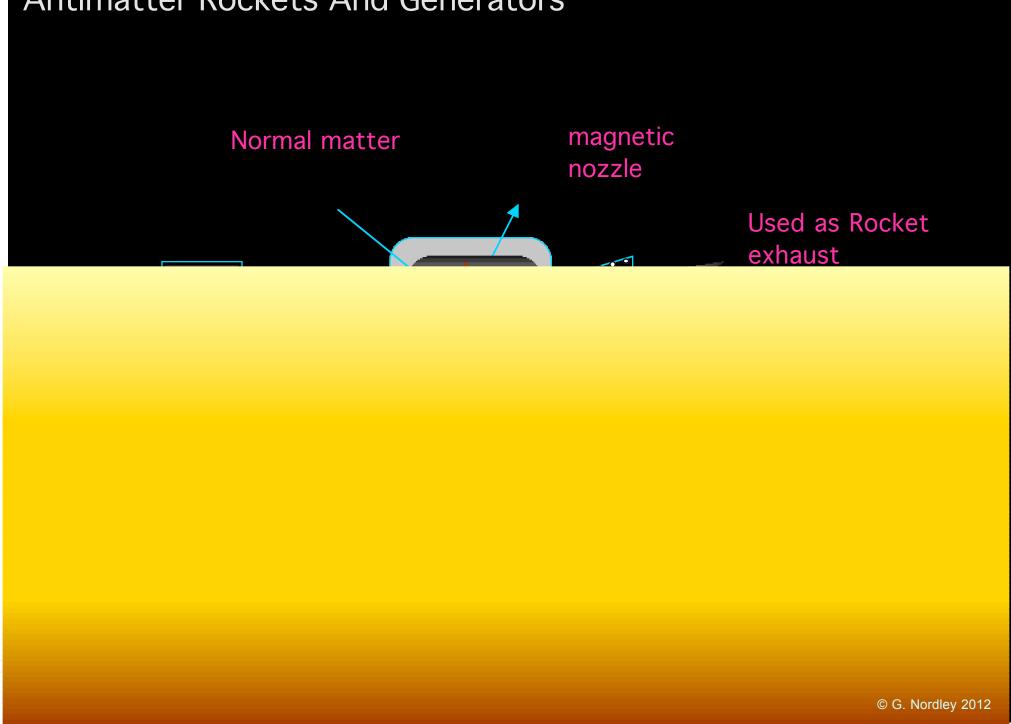
ORION (NASA

Image]

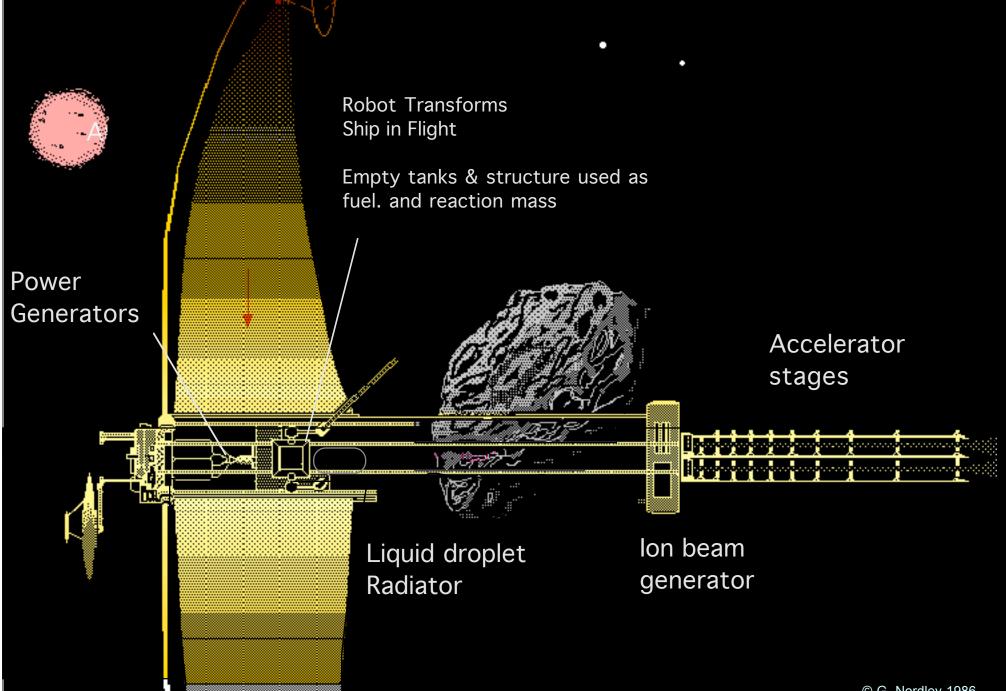
Antimatter



Antimatter Rockets And Generators



Antimatter-Powered Electric Rocket



Interstellar Ramjets

About 1 atom per cubic centimeter

Concept: Scoop up protons and atoms and fuses them.

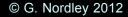
"Free" Reaction Mass and Energy

But Thrust and Acceleration are Low

To get 1 kg/s at .2c, (\approx 60 Mm/s) Scoop Radius = 1,782 km

Proton fusion is difficult.

Drag may limit scoop to a few tenths of c



Beam Supported Propulsion

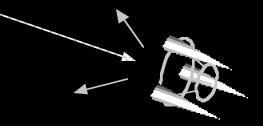
Leaves much propulsion mass at "rest"

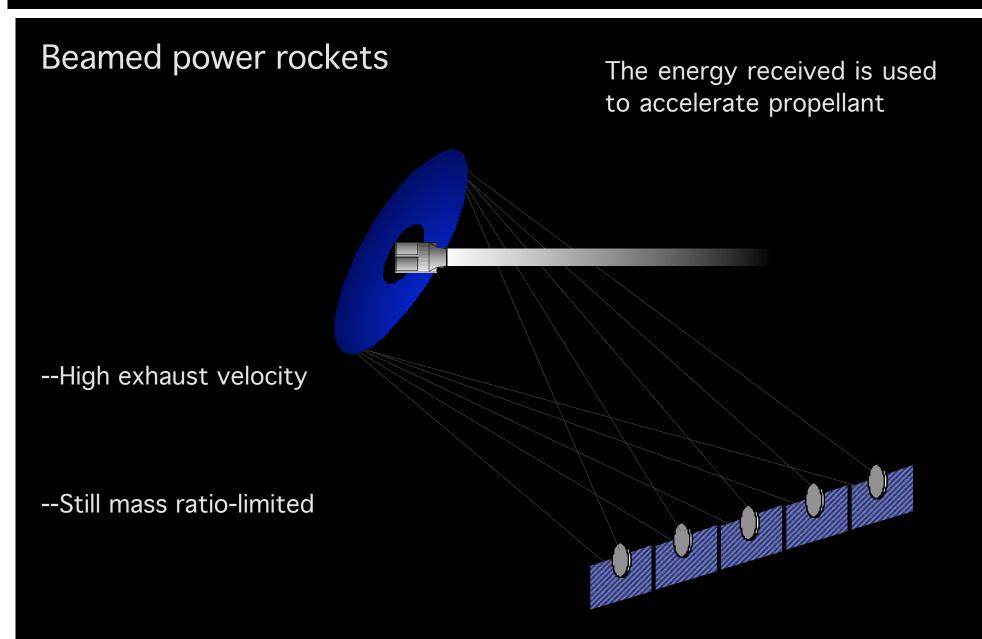
--but beam must go to spacecraft!

Beam Powered Rockets

Photon Sails: Sun, Lasers, Microwaves

Mass Beam Riders





Light or Microwaves are beamed to spacecraft

Photon Sails

PHOTONS ARE EASY TO MAKE

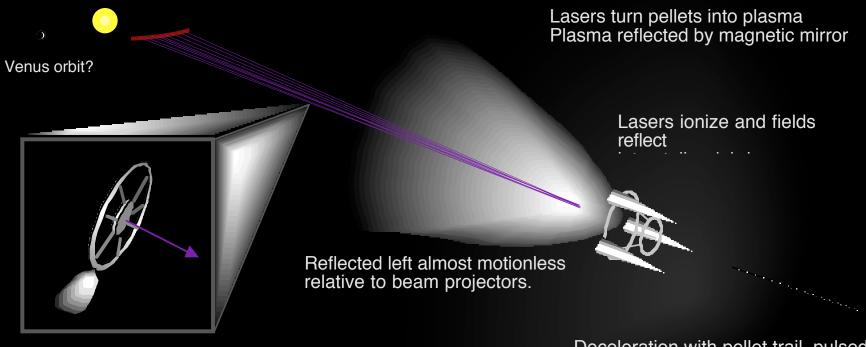
Arrays collect many times the spacecraft kinetic energy

Reflected Photons Lose Energy

Mass Beam Propelled Spacecraft

Solar power stations power mass beam drivers

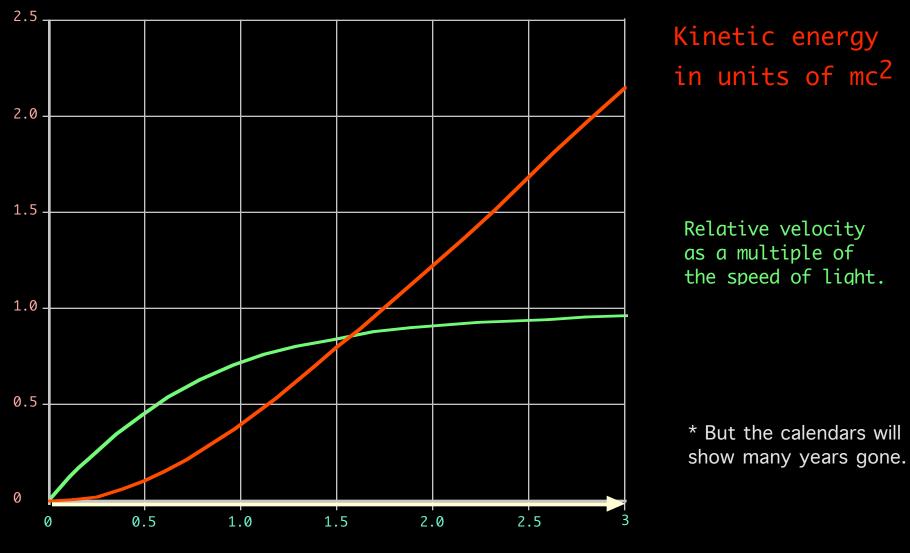
Beam drivers shoot pellets at spacecraft with the right velocity



Pellets steer themselves to beamrider

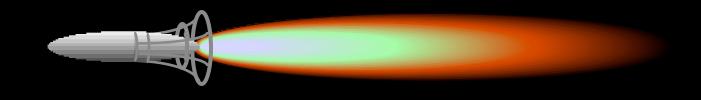
Deceleration with pellet trail, pulsed fusion, or by some other means.

RELATIVITY: NO LIGHTSPEED BARRIER FOR TRAVELERS*



Star-map velocity as a multiple of the speed of light

Life on a Starship



Power

Gravity?

Provisioning

Recycling

Cold Sleep?

Elbow Room



ISS NASA Image

REGENERATIVE LIFE SUPPORT - INPUTS AND OUTPUTS From: Space Station Freedom Design Parameters -NASA SSP 30362 (1990).

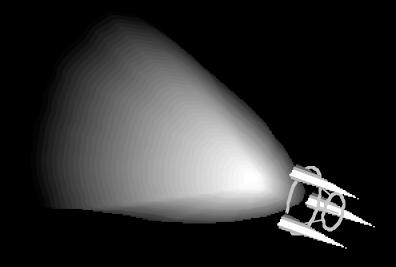
INPUTS	kg/person/day
Oxygen	0.83
Dry Food	0.62
Water in Food	1.15
Food Preparation Water	0.79
Drinking Water	1.61
Oral Hygiene Water	0.36
Hand and Face Wash Water	1.81
Shower Water	5.44
Clothes Wash Water	12.47 *
Dish Wash Water	5.44
Urinal/Commode Flush Water	0.49
Total: 31.0 kg/person/day	

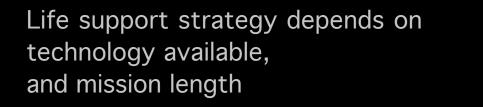
OUTPUTS - kg/person/day

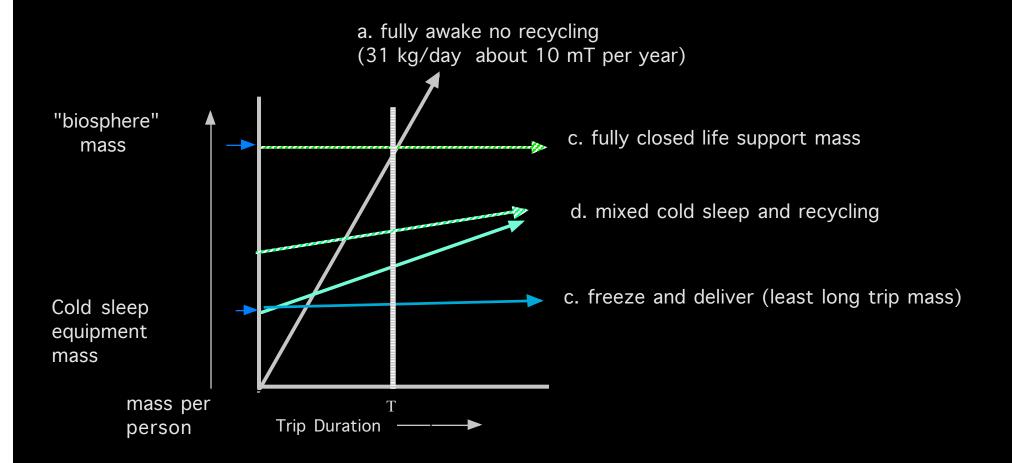
Carbon Dioxide	1.34				
Urine Solids	0.06				
Sweat Solids	0.02				
Fecal Solids	0.03				
Water vapor from skin and lungs	2.28				
Urine					
Feces Water	0.09				
Hygiene Waste Water	7.18				
Hygiene Evaporated Water	0.44				
Clothes Wash Water	11.87				
Clothes Wash Water Vapor	0.60*				
Food Preparation Water Vapor	0.04				
Dish Wash Water	5.41				
Dish Wash Evaporated Water	0.03				
Urinal and Commode Flush Water	0.49				
otal: 31.0 kg/person/day					

If we have complete water recycling, expendable supplies per person per year of active time are about

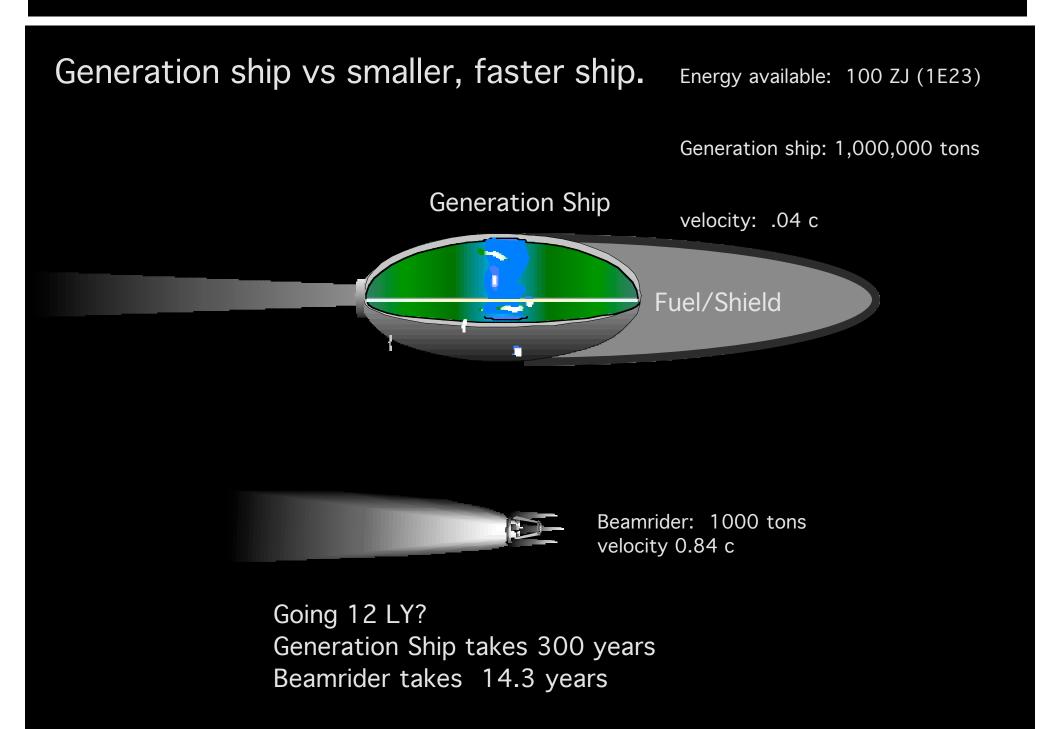
500 kg



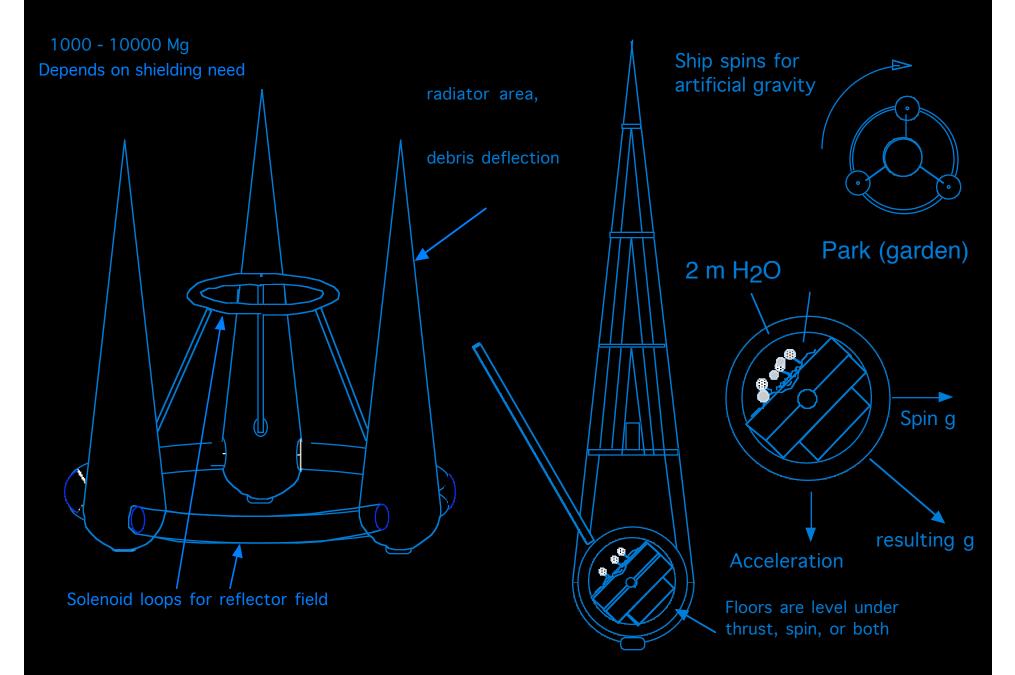




Trade between life support closure and suspended animation.



STARSHIP DESIGN STUDY



Some personal conclusions:

No "generation ships." (Lifetime extension, propulsion advances)

Starships with human crews won't be launched until relative velocities of 0.3 to 0.5 c are achieved.

Interstellar Journeys will take a few years to a few decades (Solar System time)

Robotic systems will explore and prepare systems for humans.

To Proxima Centauri with a 1000 ton ship

3 g, 0.95 reflection, 0 residual velocity

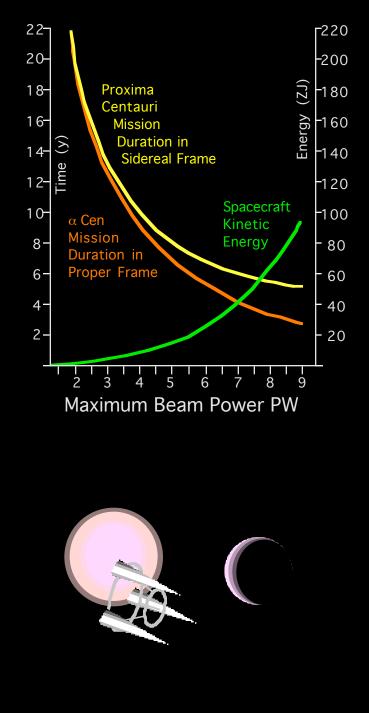
Peak power required \approx 40,000 TW

Coasting velocity = .886 C

Acceleration time: 200 d Earth-frame 152 d spaceship

Transit time ≈ 5.33 years Earth ≈ 3 years, spaceship

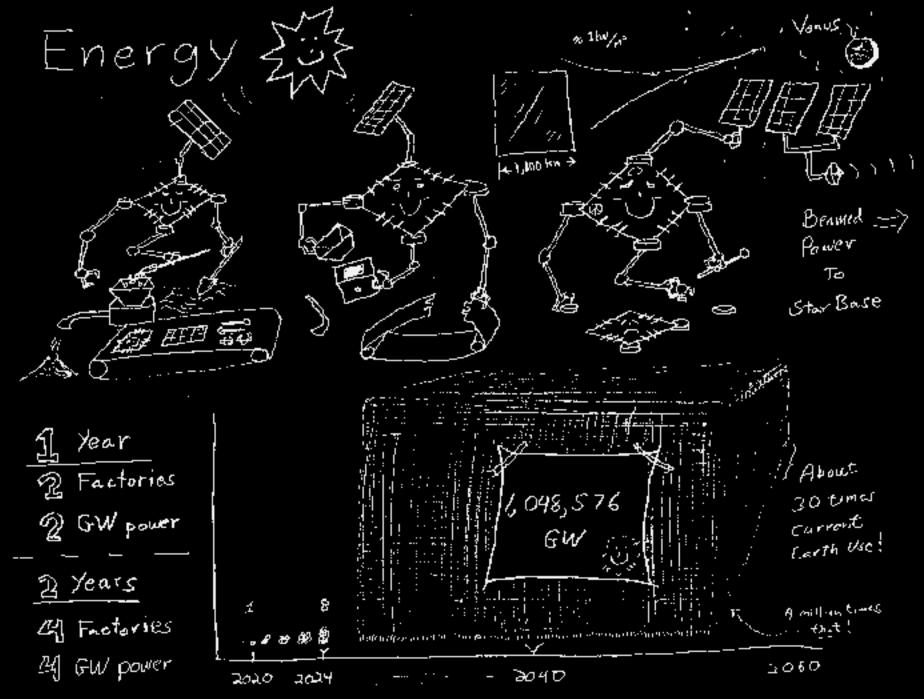
20% Efficiency, Arrays to ship KE = 565 E21 J

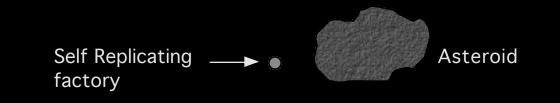


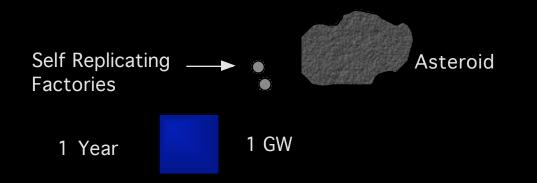
Turning Asteroids Into Solar Power Stations

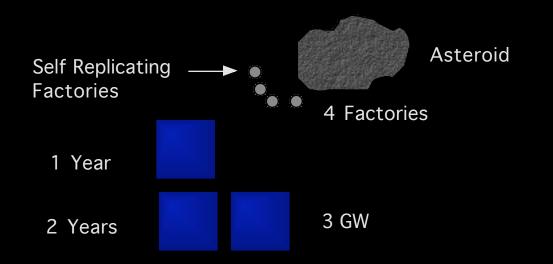


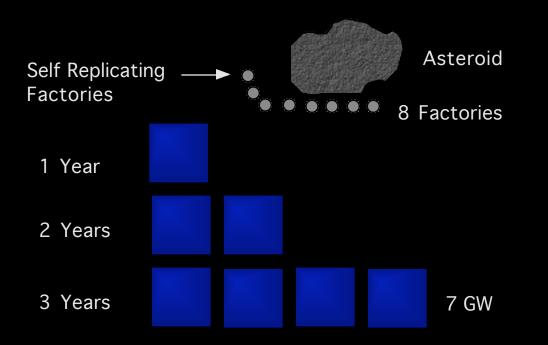
Exponential Growth

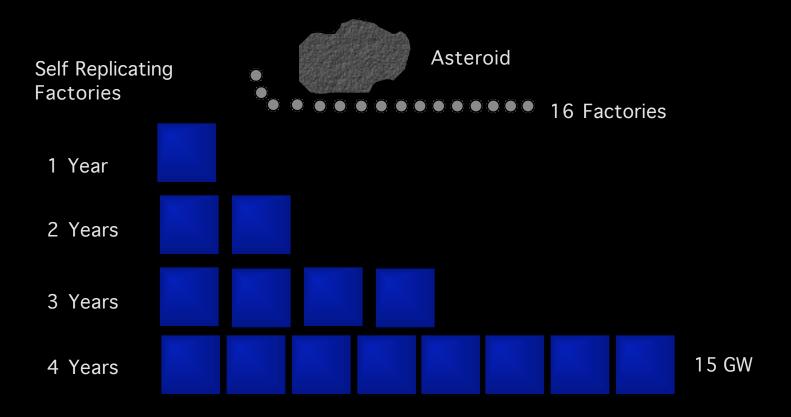


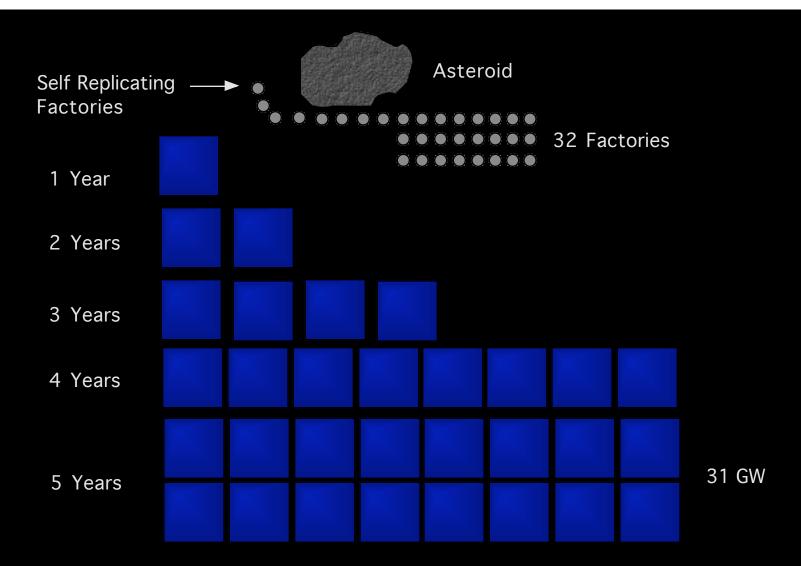


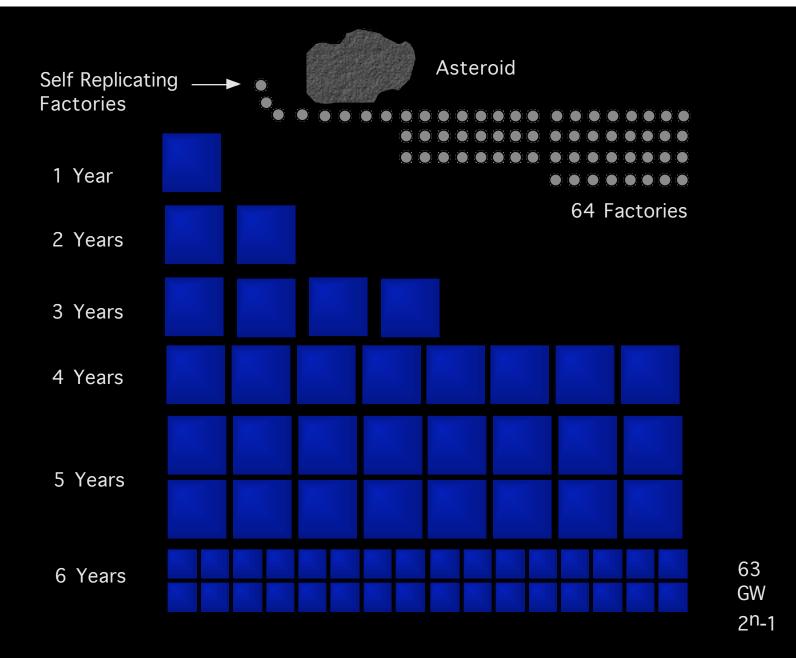


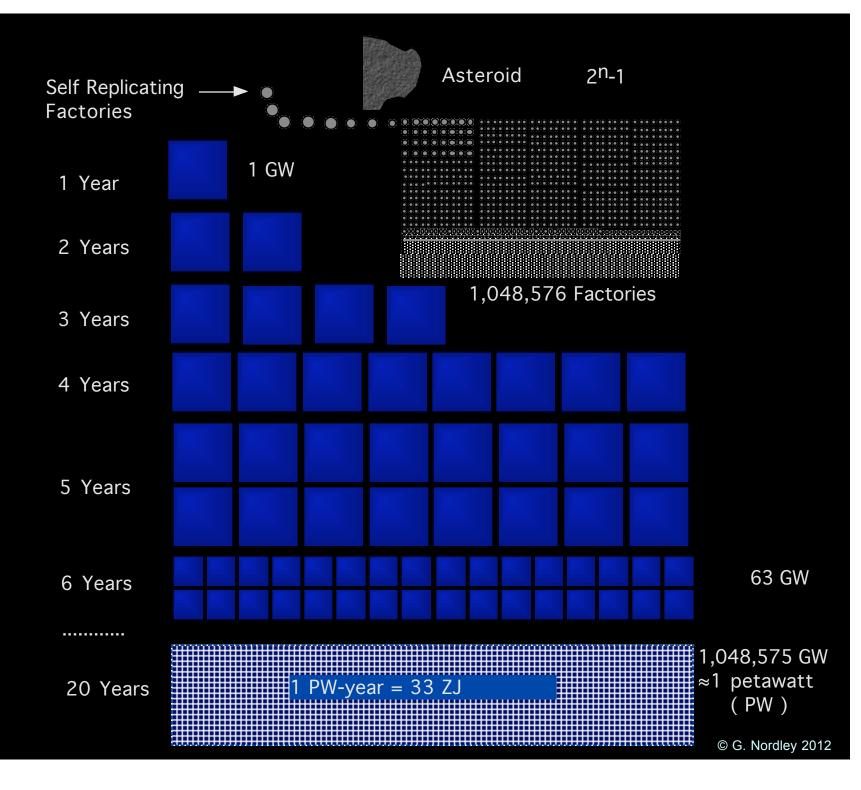


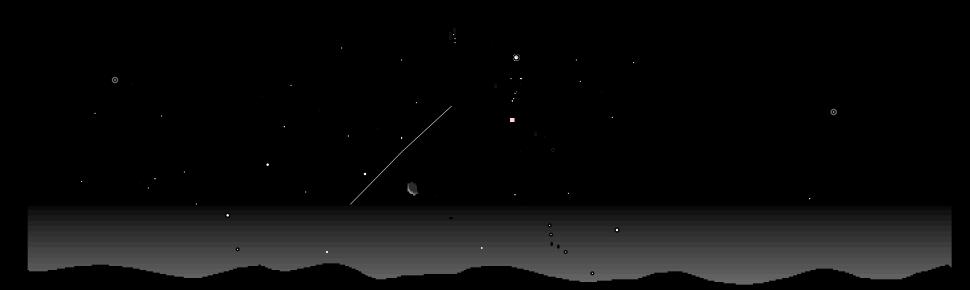












Array width: 10,000 km

Array length 55,000,000 km

Array area $55 \text{ E } 15 \text{ m}^2$

arc length 30 deg

Fraction of a Dyson sphere: 3.75 E-6

Power collected 550 E6 TW

Starships powered to 0.86c: 36,000

Needed to build: 40 Y

Radius of raw material asteroid ≈ 300 km

THOUGHTS ON INTERSTELLAR CIVILIZATION

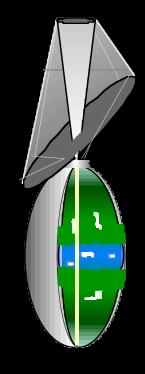
Divergent human cultures with common heritage

---Slow paced, but not totally isolated

--extended life, robot labor, large scale

Trade (if any) in unique artifacts

--robots will build what is needed from local material



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Star travelers of different intelligences

- --could form their own culture
- --biological impulses subordinate to intelligence
- --same physical laws, similar engineering achievements cultural interchange
- --so, cultural convergence among different species

This may already have happened....

Interstellar Travel Summary:

The energy requirements for human interstellar flight are huge.

Given energy, relative velocities approaching c are possible

The important technologies are robotics and space access.

Biotech advances make it easier.

Starfaring may be like Early Seafaring; long journeys seldom im port.

Ad Astra

Gerald David Nordley OryCon 2010,

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